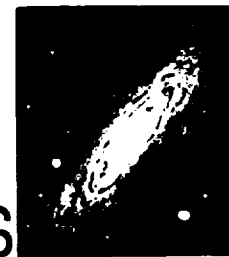
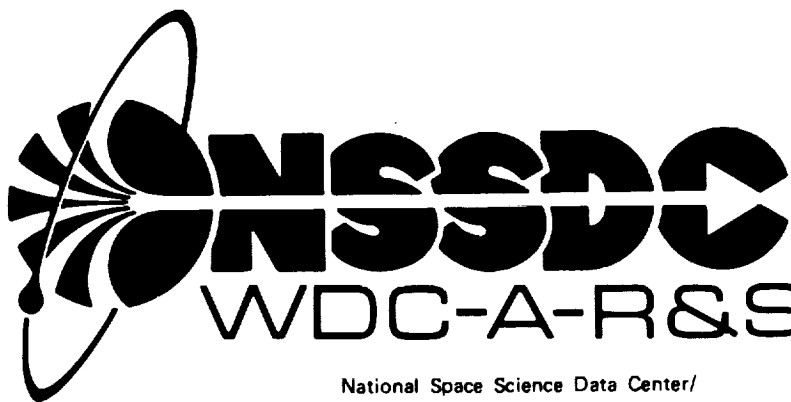


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National Space Science Data Center/
World Data Center A For Rockets and Satellites

SOUTHERN DURCHMUSTERUNG

(Schönfeld 1886)

Documentation for the Machine-Readable Version

(NASA-TM-105062) SOUTHERN DURCHMUSTERUNG
(SCHÖNFELD 1886): DOCUMENTATION FOR THE
MACHINE-READABLE VERSION (NASA) 17 p

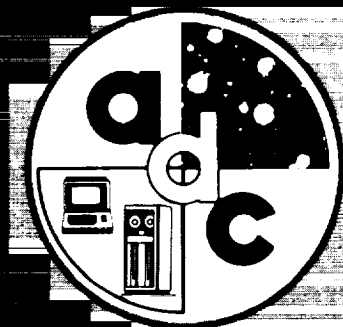
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April 1989

SOUTHERN DURCHMUSTERUNG

(Schönfeld 1886)

Documentation for the Machine-Readable Version

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April 1989

National Space Science Data Center (NSSDC)/
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Abstract

The machine-readable version of the catalog, as it is currently being distributed from the Astronomical Data Center, is described. The *Southern Durchmusterung* (SD) was computerized at the Centre de Données Astronomiques de Strasbourg and at the Astronomical Data Center at the National Space Science Data Center, NASA/ Goddard Space Flight Center. Corrigenda listed in the original SD volume and published by Küstner and Sticker have been incorporated into the machine file. In addition, one star indicated to be "missing" in a published list, and later verified, is flagged so that it can be omitted from computer plotted charts if desired. Stars deleted in the various errata lists have been similarly flagged, while those with revised data are flagged and listed in a separate table. This catalog covers the zones -02° to -23° ; zones $+89^{\circ}$ to -01° (the *Bonner Durchmusterung*) are included in a separate catalog available in machine-readable form.

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1.0 Introduction

1.1 Description

The *Southern Durchmusterung* (SD, Schönfeld 1886, Becker 1949, Schmidt 1967) is a visual survey of stars in the declination zones -02° to -23° , completed as an extension to Argelander's (1859-62) monumental *Bonner Durchmusterung* (BD). Schönfeld's survey was carried out using the same methods as had been used for the BD, which Schönfeld had helped to compile as one of Argelander's assistants. The procedure consisted of allowing the telescope to drift along the mean declination of each zone and recording the positions and magnitudes of stars crossing the transit line of each field. The goal of the survey was to extend the BD to declination -23° (a plan originally adopted by Argelander) with approximately the same magnitude limits, although the primary instrument was of larger aperture (159 mm) than the 78-mm telescope used for the BD. Thus, whereas the BD magnitude estimates extend to $9^m.4$ with all fainter stars assigned a magnitude of 9.5, the SD magnitude estimates extend to $9^m.9$ with all fainter stars assigned a magnitude of 10. The fainter limit of the SD results in an increase in the average number of stars in each zone of a thousand or more over the BD. As with the BD, the SD contains a rather large number of stars fainter than $10^m.0$ and even occasionally as faint as 11^m . Positions are given to the nearest $0^s.1$ in right ascension and $0'.1$ in declination as in the BD.

This documentation is intended to fully describe the machine-readable version of the SD. It includes detailed descriptions of the format and the procedure by which the computer file was created. Lists of all corrections made to the original data as a result of published corrigenda, stars deleted according to overstriking in the printed catalogs or their inclusion in lists of "missing" stars, and stars inserted in later editions are given in a separate table. Zone statistics for the catalog are also given in a table. No other corrections or changes have been incorporated into the original data, e.g., from more modern positions and magnitudes or from comparison with other catalogs. The document is intended to enable users to process the data without problems, guesswork, or further literature consultation. For more detailed descriptions of how the observations were made and for additional statistics of star counts and distributions within each zone, the source references should be consulted. The reader can also find additional information on the compilation and scale of the SD, published in English, in papers by Pickering (1885, see pages 361 and 478; 1892; 1913). A copy of this document should accompany any machine version of the catalog originating from the Astronomical Data Center or from the Center de Données Astronomiques de Strasbourg.

1.2 Source References

- Becker, F. 1949, *Bonner Durchmusterung*, Südlicher Teil, Deklinations-Zonen -2° bis -22° Sternverzeichnis, zweite, berichtigte Auflage (Bonn: Ferd. Dümmlers Verlag).
- Schmidt, H. 1967, *Bonner Durchmusterung*, Südlicher Teil, Deklinations-Zonen -2° bis -22° Sternverzeichnis, dritte Auflage (Bonn: Ferd. Dümmlers Verlag).
- Schönfeld, E. 1886, *Bonner Sternverzeichniss*, vierte Sektion, *Astronomische Beobachtungen auf der Sternwarte der Königlichen Rheinischen Friedrich-Wilhelms-Universität zu Bonn* 8, Part IV (Bonn: Adolph Marcus).

2.0 Structure

2.1 File Summary

The machine version of the *Southern Durchmusterung* consists of a single file. Table 1 gives the machine-independent file attributes. All logical records are of fixed length; if the catalog is received on magnetic tape, it will contain blocks of fixed length (as noted below) except that the last block of each file may be short.

<i>Southern Durchmusterung</i> (Schönfeld 1886)				
File	Contents	Record Format	Logical Record Length	Total Number of Logical Records
1	Data	FB	32	134834

Table 1. Summary Description of Catalog Files: FB = Fixed length blocks (last may be short)

The information contained in the above table is sufficient for a user to describe the indigenous characteristics of the machine-readable version of the *Southern Durchmusterung* to a computer. Information easily varied from installation to installation, such as block size (physical record length), blocking factor (number of logical records per physical record), total number of blocks, density, number of tracks, and character coding (ASCII, EBCDIC) for tapes, is not included but should always accompany secondary copies if any are supplied to other users or installations.

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2.2 Catalog (File 1 of 1)

Table 2 gives a byte-by-byte description of the contents of the data file. A suggested Fortran format specification for reading each data field is included and can be modified depending upon individual programming and processing requirements (Fortran 77 character string-type formats are used); however, caution is advised when substituting format specifications, since some of the data fields contain character data. Default (null) values are always blanks in data fields for which primary suggested formats are given as A. Null values are also not specified for numerical fields that always contain valid data.

Byte(s)	Units	Suggested Format	Default Value	Data
1-2	---	A2	---	Catalog prefix (SD)
3-5	---	I3	---	Zone
6-10	---	I5	---	Star number
11	---	A1	---	Code
12-15	mag	F4.1	---	Visual magnitude
16-17	hours	I2	---	Right ascension, α
18-19	min	I2	---	α
20-23	sec	F4.1	---	α
24	---	A1	---	Sign of declination, δ
25-26	°	I2	---	δ
27-32	'	F6.3	---	δ

Table 2. Data File Record Format

Catalog prefix	The letters "SD" are included in each record in order to distinguish the SD from all other machine-readable Durchmusterungen. Thus, all of the DM catalogs can conceivably be combined into a single file, and the individual source catalogs will still be identifiable.
Zone	The δ zone part of the SD number. The sign is always in byte 3, with preceding zeros on single-digit zones where appropriate.
Star number	Sequentially increasing star number within the specified zone.
Code	Upper and lower case codes. All lower case letters represent supplemental (footnoted) stars added to the catalog via published corrigenda. Upper case letters and asterisks are flags that indicate changes to the original data or to the status of a star in the catalog. They have the following meanings: <ul style="list-style-type: none"> * Data have been corrected as a result of corrigenda, or there are special notes associated with the star. All changes are given in "Appendix A. Changes to Catalog Data", Table 4 on page 11, with appropriate notes. D The star has been deleted in a later edition of the catalog. This was done by overstriking entries with horizontal lines. M The star was noted as "missing" in a list published by Pickering (1907). This entry (there is only one such star) was verified by R. A. Downes and at the ADC.

Visual magnitude

Magnitude as estimated by the observer or magnitude code to denote non-numerical entries in the published catalogs. The following codes are used (we describe all codes used in the *Durchmusterung* catalogs; not all are used in the SD):

20.0 *neb* (denoting a nebula);

30.0 *var* (denoting variability);

40.0 *nova* or *nova?*;

50.0 *cum* (denoting the cumulative [integrated] magnitude estimate of a cluster of stars).

Equatorial coordinates

Equinox 1855. For δ^m in the SD, only bytes 27-30 are used, so the data can be read with format F4.1. Bytes 31-32 were reserved to maintain a uniform format for all DM catalogs, since they are used in the southern zones of the *Cape Photographic Durchmusterung*.

3.0 History

3.1 Remarks

The data in the machine-readable *Southern Durchmusterung* were keyed directly to disk storage from the published catalogs at the Centre de Données Astronomiques de Strasbourg and at the Astronomical Data Center of the National Space Science Data Center at the NASA Goddard Space Flight Center. Progress reports on the keypunching of the *Durchmusterung* catalogs have been published over the years in the *Bulletin d'Information du Centre de Données Stellaires* (Couteau *et al.* 1983; Wagner 1984, 1986; Warren 1987). The distribution of SD work is given in Table 3.

Zones	Location
-01°	Astronomical Data Center/NSSDC
-02° to -21°	Centre de Données Astronomiques de Strasbourg (CDS)
-22° to -23°	Astronomical Data Center/NSSDC

Table 3. Distribution of Computerization Work for SD Data

Most of the data entry work was done at the CDS, where the data were also verified. The ADC zones -01° and -23°, although not often used for stellar designations (the BD is used in zone -01° and the *Córdoba Durchmusterung* in zone -23°), were added for completeness. The three ADC zones were computerized and proofread by the first author (WHW). All zones were carefully examined, checked for sequencing and record counts, reformatted, and merged in the correct order at the ADC, where the final catalog was assembled. The published corrigenda lists were cross checked against the reprinted editions (1949, 1967) of the catalog. The final catalog was run through a verification program that checked numerical sequencing of the SD numbers, monotonic increase in right ascension, and allowed data ranges. All cases where stars are out of RA order were checked in the original data to verify that their positions are as in the original catalog or the corrigenda.

The final file is ordered north to south strictly by SD number, *i.e.*, in the zone order -01°, -02°, ..., -23°. Users should note, however, that all stars are not strictly in right ascension order within each zone. This is because individual stars are occasionally out of RA order in the original catalog and because of corrections inserted from the corrigenda. Thus, if the machine catalog is sorted by increasing RA, *e.g.*, for search purposes, some SD numbers will become disordered.

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4.0 Acknowledgments and References

4.1 Acknowledgments

The authors express appreciation to M. J. Wagner, M. Maslo, and R. Bonnet, who keyed the SD data to disk at the CDS using preprocessing software written and implemented by F. Ochsenbein. We thank Dr. R. A. Downes for bringing the cases of "missing" BD and SD stars to our attention. We are grateful for the support of the CDS director, Dr. C. Jaschek, during the course of the work.

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Warren, W. H. Jr. 1987, *Inf. Bull. CDS* No. 32, p. 67.

Appendix A. Changes to Catalog Data

As mentioned previously, all changes made to SD data and known to us have been incorporated into the present machine-readable version. Unlike the *Bonner Durchmusterung*, where there had been numerous stars added and deleted, changes in supplemental entries from version to version, and a large number of data corrections, the changes to the SD are relatively few and can be documented in a single table. There were two stars added as supplementary entries ("a" suffix), two stars deleted, and one star published as "missing" by Pickering (1907) and subsequently verified by R. A. Downes of Applied Research Corporation. A small list of corrections was also included in the introduction to the original published catalog. These modifications to the printed catalog are flagged by codes in byte 11 of the affected records and are listed in Table 4 below.

Star	Datum	For	Read	Remarks
-06 5412	δ	28.8	28.3	
-07 5419	δ	43.0	53.0	
-09 553a				Added by Küstner (1918).
-09 743	δ	45.2	54.2	
-09 4403	m_v	8.9	9.3	
-10 4102a				Added by Küstner (1918).
-12 404	δ	12.6	42.6	Sticker (1935).
-14 133	δ	32.1	33.1	
-14 5658	δ	46.1	45.1	
-15 1427	δ	26.9	25.9	"Missing" star (Pickering 1907) flagged with "M".
-15 4676				
-16 4872	δ	50.4	54.0	
-18 4706	δ	52.8	53.2	
-19 270	δ	38.9	34.5	
-19 6001	α^s	32.0	34.4	
-20 475	δ	2.0	6.9	Deleted by Küstner (1918). Puts out of order with 3949. Deleted by Küstner (1925).
-20 477	δ	54.2	58.6	
-20 1041				
-20 1697	α^s	23.6	26.4	
-20 1697	δ	28.3	28.0	
-20 2161	δ	28.4	23.8	
-20 3648	α^s	39.0	46.2	
-20 3950	α^s	48.5	45.3	
-20 4293a				
-20 5349	α^s	11.3	9.6	
-20 5671	α^s	26.4	27.2	
-21 957	m_v	8.9	8.8	
-22 3885	α^s	14.1	11.3	
-22 5735	α^s	45.5	55.5	

Table 4. Changes Made to SD Catalog Data

A.1 Zone Statistics

The changes that have been made to the SD since the original work of Schönfeld have resulted in a situation that can be very confusing to a catalog user who wishes to verify that disagreements between the published catalog and the machine version are valid. For example, one usually checks the number of logical records in a sequentially numbered catalog with the highest sequential number in order to verify that no objects are missing or that there are no duplicate records, etc. Because of the supplemental SD stars, this is not possible in some zones of the SD nor for the total.

It is also useful to be able to determine how many objects have been added, deleted, etc. without having to process the whole catalog. Therefore, we summarize the zone statistics for the SD in Table 5 for the convenience of users of this new machine version. Note that, for some zones, the highest SD number does not determine the star count, but, since records for deleted stars have been flagged rather than removed, the number of deleted stars has no influence on the record count. The "Stars Added" column shows how many stars have been added to the catalog as a result of the corrigenda, while the "Stars Deleted" column provides a count of the number of stars flagged with a "D" in each zone. The number of supplemental stars for each zone is the total (1886 plus corrigenda), hence the highest star number added to the number of supplemental stars should produce the record count. The "M" star ($-15^{\circ} 4676$) is still counted in the last column.

Zone	Number of Records	Highest Star Number	Stars Added	Stars Deleted	Stars Missing	Supplemental Stars	Number of "Active" Stars
-01	692	692	-	-	-	-	692
-02	6099	6099	-	-	-	-	6099
-03	5763	5763	-	-	-	-	5763
-04	6026	6026	-	-	-	-	6026
-05	6117	6117	-	-	-	-	6117
-06	6365	6365	-	-	-	-	6365
-07	6153	6153	-	-	-	-	6153
-08	6247	6247	-	-	-	-	6247
-09	6327	6326	1	-	-	1	6327
-10	6231	6230	1	-	-	1	6231
-11	6206	6206	-	-	-	-	6206
-12	6617	6617	-	-	-	-	6617
-13	6521	6521	-	-	-	-	6521
-14	6621	6621	-	-	-	-	6621
-15	6545	6545	-	-	1	-	6545
-16	6419	6419	-	-	-	-	6419
-17	6878	6878	-	-	-	-	6878
-18	6429	6429	-	-	-	-	6429
-19	6570	6570	-	-	-	-	6570
-20	6724	6724	-	2	-	-	6722
-21	6538	6538	-	-	-	-	6538
-22	6265	6265	-	-	-	-	6265
-23	481	481	-	-	-	-	481
Totals	134834	134832	2	2	1	2	134832

Table 5. Zone Statistics for the SD

Appendix B. Sample Listing

The sample listing given on the following pages shows logical records exactly as they are recorded in the machine-readable version of the catalog. Groups of records from the beginning and end of the file are illustrated. The beginning of each record and the bytes within the record are indicated by the column heading index across the top of each page (digits read vertically).

